# Optimum Impulsive Hohmann Coplanar Elliptic Transfer <br> Using Energy Change Concept Part II 

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#### Abstract

We present an elementary approach for the optimization of the elliptic coplanar coaxial Hohmann type transfer arising from the first principles. We assign the minimized increments of velocities at peri-apse and apo-apse by equating to zero the gradient of $\Delta_{v 1}+\Delta_{v 2}$, then resolving a second degree algebraic equation in the variable x (the ratio of the velocities before and after the initial impulse). We consider the four feasible configurations, and we assign the most economic one. By setting $e_{1}=0, e_{2}=0$ for the terminal orbits, we confront the original circular Hohmann transfer case promptly.


Keywords: Rocket dynamics, elliptic Hohmann transfer, optimization

## 1. Background

Orbit transfer is a major subject with regard to placing a spacecraft in an orbit around the Earth. The velocity increments are directly proportional to motor system thrusts of the rocket space vehicle. Consequently it is proportional to propellant fuel consumption. It is most convenient to regard the transfer problem as a problem of change of energy [1]. We utilize well known geometric properties of conic sections and ordinary calculus. The main two types of orbit transfer are the Hohmann and the Bi-Elliptic. For each type we face the coplanar and the non coplanar cases. The criterion for optimality is the minimization of the characteristic velocity for the maneuver [2], [3]. The literature of optimal transfer is and so on extensive, we may recall the works by Prussing [3], Palmore [4], Edelbaum [5], Barrar [6], Marec [7], Lawden [8], Hiller [9] and Altman and Pistiner [10]. It is established


Figure 1 Apogee of transfer orbit coincides with apogee of final orbit (Initial impulse at perigee)
that the minimum total velocity increment solutions for specified trajectory end point conditions, are attainable directly by methods of the differential calculus.

## 2. Method and Results

We begin by the first configuration for a two impulse Hohmann elliptic transfer of a space vehicle (Fig. 1). We consider the following relationships:

$$
\begin{align*}
& I_{1}=\Delta v_{1}=v_{A 2}-v_{A 1}=x v_{A 1}-v_{A 1}=(x-1) v_{A 1}  \tag{1}\\
& I_{2}=\Delta v_{2}=v_{B 2}-v_{B 1}  \tag{2}\\
& \quad v_{A 2}=\sqrt{\frac{\mu\left(1+e_{T}\right)}{a_{T}\left(1-e_{T}\right)}} \quad v_{A 1}=\sqrt{\frac{\mu\left(1+e_{1}\right)}{a_{1}\left(1-e_{1}\right)}} \\
& \quad v_{B 2}=\sqrt{\frac{\mu\left(1-e_{2}\right)}{a_{2}\left(1+e_{2}\right)}} \quad v_{B 1}=\sqrt{\frac{\mu\left(1-e_{T}\right)}{a_{T}\left(1+e_{T}\right)}} \tag{3}
\end{align*}
$$

where

$$
\begin{gather*}
x=\frac{x v_{A 1}}{v_{A 1}}=\frac{\text { velocityafterperi }- \text {-apseinitialimpulse }}{\text { velocitybeforeperi }- \text {-apseinitilimpulse }} \quad x>1 \\
x=\sqrt{\frac{\frac{\mu\left(1+e_{T}\right)}{a_{T}\left(1-e_{T}\right)}}{\frac{\mu\left(1+e_{1}\right)}{a_{1}\left(1-e_{1}\right)}}} \tag{4}
\end{gather*}
$$

From the geometry of Fig. 1, we have

$$
\begin{align*}
& a_{T}\left(1+e_{T}\right)=a_{2}\left(1+e_{2}\right)  \tag{5}\\
& a_{T}\left(1-e_{T}\right)=a_{1}\left(1-e_{1}\right) \tag{6}
\end{align*}
$$

From Eqs. (5), (6), we get

$$
\begin{equation*}
\frac{1-e_{T}}{1+e_{T}}=\frac{a_{1}\left(1-e_{1}\right)}{a_{2}\left(1+e_{2}\right)} \tag{7}
\end{equation*}
$$

whence, from Eqs. (2), (7)

$$
\begin{equation*}
\Delta v_{2}=\sqrt{\frac{\mu\left(1-e_{2}\right)}{a_{2}\left(1+e_{2}\right)}}-\sqrt{\frac{\mu a_{1}\left(1-e_{1}\right)}{a_{T} a_{2}\left(1+e_{2}\right)}} \tag{8}
\end{equation*}
$$

From Eqs. (4), (7), we acquire

$$
\begin{equation*}
x=\sqrt{\frac{1+e_{T}}{1+e_{1}}}>1 \tag{9}
\end{equation*}
$$

i.e.

$$
e_{T}=x^{2}\left(1+e_{1}\right)-1
$$

From Eqs. (5), (6), we find

$$
a_{T}=\frac{a_{1}\left(1-e_{1}\right)}{1-e_{T}}=\frac{a_{2}\left(1+e_{2}\right)}{1+e_{T}}
$$

Whence

$$
\begin{equation*}
a_{T}=\frac{a_{1}\left(1-e_{1}\right)}{2-x^{2}\left(1+e_{1}\right)}=\frac{a_{2}\left(1+e_{2}\right)}{x^{2}\left(1+e_{1}\right)} \tag{10}
\end{equation*}
$$

We can easily derive

$$
v_{B 1}=\sqrt{\frac{\mu\left\{2-x^{2}\left(1+e_{1}\right)\right\}}{a_{2}\left(1+e_{2}\right)}}
$$

Therefore

$$
\begin{gather*}
\Delta v_{2}=\sqrt{\frac{\mu\left(1-e_{2}\right)}{a_{2}\left(1+e_{2}\right)}}-\sqrt{\frac{\mu\left\{2-x^{2}\left(1+e_{1}\right)\right\}}{a_{2}\left(1+e_{2}\right)}}  \tag{11}\\
\Delta v_{1}=\sqrt{\frac{\mu\left(1+e_{1}\right)}{a_{1}\left(1-e_{1}\right)}}(x-1)
\end{gather*}
$$

For the optimum condition:

$$
\begin{equation*}
\frac{d}{d x}\left(\Delta v_{T}\right)=\frac{d}{d x}\left(\Delta v_{1}\right)+\frac{d}{d x}\left(\Delta v_{2}\right)=0 \tag{12}
\end{equation*}
$$

Let

$$
\begin{array}{ll}
b_{1}=a_{1}\left(1-e_{1}\right) & b_{2}=a_{1}\left(1+e_{1}\right) \\
b_{3}=a_{2}\left(1-e_{2}\right) & b_{4}=a_{2}\left(1+e_{2}\right) \tag{13}
\end{array}
$$

Whence by differentiation w.r.t. the variable $x$

$$
\begin{align*}
& \frac{d}{d x}\left(\Delta v_{1}\right)=\sqrt{\frac{\mu\left(1+e_{1}\right)}{b_{1}}}=v_{A 1}  \tag{14}\\
& \frac{d}{d x}\left(\Delta v_{2}\right)=\sqrt{\frac{\mu}{b_{4}}} \frac{x\left(1+e_{1}\right)}{\sqrt{2-x^{2}\left(1+e_{1}\right)}}  \tag{15}\\
& \text { i.e. } \\
& \sqrt{\frac{\mu\left(1+e_{1}\right)}{b_{1}}}+\sqrt{\frac{\mu}{b_{4}}} \frac{x\left(1+e_{1}\right)}{\sqrt{2-x^{2}\left(1+e_{1}\right)}}=0 \tag{16}
\end{align*}
$$

After some reductions and rearrangements, we get

$$
\begin{equation*}
(x)_{M i n}= \pm \sqrt{\frac{2 b_{4}}{\left(b_{1}+b_{4}\right)\left(1+e_{1}\right)}} \tag{17}
\end{equation*}
$$

Or in explicit form

$$
\begin{equation*}
(x)_{M i n}=\sqrt{\frac{2 a_{2}\left(1+e_{2}\right)}{\left(1+e_{1}\right)\left\{a_{1}\left(1-e_{1}\right)+a_{2}\left(1+e_{2}\right)\right\}}} \tag{18}
\end{equation*}
$$

By substitution in Eqs. (9), (10) for the value of $(x)_{\text {Min. }}$. Eq. (18), we get the unique values for ( $\mathrm{a}_{T}, \mathrm{e}_{T}$ ), namely

$$
\begin{align*}
& \left(a_{T}\right)_{M i n}=\frac{1}{2}\left[a_{1}\left(1-e_{1}\right)+a_{2}\left(1+e_{2}\right)\right]  \tag{19}\\
& \left(e_{T}\right)_{M i n}=\frac{-a_{1}\left(1-e_{1}\right)+a_{2}\left(1+e_{2}\right)}{a_{1}\left(1-e_{1}\right)+a_{2}\left(1+e_{2}\right)} \tag{20}
\end{align*}
$$

Which shows that the generalized Hohmann transfer is itself a minimum transfer system.

Now we evaluate the minimum characteristic velocity $\left(\Delta v_{T}=\Delta v_{1}+\Delta v_{2}\right)_{M i n}$, we have

$$
\begin{equation*}
\Delta v_{T}=\sqrt{\frac{\mu\left(1+e_{1}\right)}{b_{1}}}(x-1)+\sqrt{\frac{\mu\left(1-e_{2}\right)}{b_{4}}}-\sqrt{\frac{\mu}{b_{4}}\left\{2-x^{2}\left(1+e_{1}\right)\right\}} \tag{21}
\end{equation*}
$$

By substitution for $x=(x)_{M i n}$, we find that

$$
\begin{equation*}
\left(\Delta v_{T}\right)_{M i n}=\sqrt{\frac{2 \mu b_{4}}{b_{1}\left(b_{1}+b_{4}\right)}}-\sqrt{\frac{2 \mu b_{1}}{b_{4}\left(b_{1}+b_{4}\right)}}+\sqrt{\frac{\mu\left(1-e_{2}\right)}{b_{4}}}-\sqrt{\frac{\mu\left(1+e_{1}\right)}{b_{1}}} \tag{22}
\end{equation*}
$$

In terms of the b's or explicitly in terms of the elements a, e

$$
\begin{align*}
& \left(\Delta v_{T}\right)_{M i n}=\sqrt{\frac{2 \mu a_{2}\left(1+e_{2}\right)}{a_{1}\left(1-e_{1}\right)\left\{a_{1}\left(1-e_{1}\right)+a_{2}\left(1+e_{2}\right)\right\}}}-\sqrt{\frac{\mu\left(1+e_{1}\right)}{a_{1}\left(1-e_{1}\right)}} \\
& +\sqrt{\frac{\mu\left(1-e_{2}\right)}{a_{2}\left(1+e_{2}\right)}}-\sqrt{\frac{2 \mu a_{1}\left(1-e_{1}\right)}{a_{2}\left(1+e_{2}\right)\left\{a_{1}\left(1-e_{1}\right)+a_{2}\left(1+e_{2}\right)\right\}}} \tag{23}
\end{align*}
$$



Figure 2 Apogee of transfer orbit coincides with perigee of final orbit (Initial impulse at perigee)

For the classical circular Hohmann transfer $e_{1}=0$ and $e_{2}=0$, whence we acquire the quite symmetric formula

$$
\begin{equation*}
\left(\Delta v_{T}\right)_{M i n}=\sqrt{\frac{\mu}{a_{1}}}\left\{\sqrt{\frac{2 a_{2}}{a_{1}+a_{2}}}-1\right\}+\sqrt{\frac{\mu}{a_{2}}}\left\{1-\sqrt{\frac{2 a_{1}}{a_{1}+a_{2}}}\right\} \tag{24}
\end{equation*}
$$

For the second configuration (Fig. 2), we have the following relationships:

$$
\begin{aligned}
& a_{1}\left(1-e_{1}\right)=a_{T}\left(1-e_{T}\right) \\
& a_{T}\left(1+e_{T}\right)=a_{2}\left(1-e_{2}\right) \\
& v_{A 1}=\sqrt{\frac{\mu\left(1+e_{1}\right)}{a_{1}\left(1-e_{1}\right)}} \\
& v_{A 2}=\sqrt{\frac{\mu\left(1+e_{T}\right)}{a_{T}\left(1-e_{T}\right)}}
\end{aligned}
$$

Let

$$
x=\frac{v_{A 2}}{v_{A 1}}=\sqrt{\frac{1+e_{T}}{1+e_{1}}}>1
$$

i.e.

$$
\begin{align*}
& e_{T}=x^{2}\left(1+e_{1}\right)-1 \\
& a_{T}=\frac{a_{1}\left(1-e_{1}\right)}{1-e_{T}}=\frac{a_{2}\left(1-e_{2}\right)}{1+e_{T}} \tag{25}
\end{align*}
$$

From Eq. (13),

$$
a_{T}=\frac{b_{1}}{1-e_{T}}=\frac{b_{3}}{1+e_{T}}
$$

Or

$$
a_{T}=\frac{b_{1}}{2-x^{2}\left(1+e_{1}\right)}=\frac{b_{3}}{x^{2}\left(1+e_{1}\right)}
$$

Now

$$
\begin{align*}
& \Delta v_{A}=v_{A 2}-v_{A 1}=x v_{A 1}-v_{A 1}=(x-1) v_{A 1} \\
& \Delta v_{A}=(x-1) \sqrt{\frac{\mu\left(1+e_{1}\right)}{b_{1}}}  \tag{26}\\
& v_{B 1}=\sqrt{\frac{\mu\left(1-e_{T}\right)}{a_{T}\left(1+e_{T}\right)}} \\
& v_{B 2}=\sqrt{\frac{\mu\left(1+e_{2}\right)}{a_{2}\left(1-e_{2}\right)}} \\
& \Delta v_{B}=v_{B 2}-v_{B 1}=\sqrt{\frac{\mu\left(1+e_{2}\right)}{b_{3}}}-\sqrt{\frac{\mu\left\{2-x^{2}\left(1+e_{1}\right)\right\}}{b_{3}}} \tag{27}
\end{align*}
$$

Optimum condition is:

$$
\frac{d}{d x}\left(\Delta v_{T}\right)=\frac{d}{d x}\left(\Delta v_{A}\right)+\frac{d}{d x}\left(\Delta v_{B}\right)=0
$$

Then, from Eqs. (26), (27) we get

$$
\sqrt{\frac{\mu\left(1+e_{1}\right)}{b_{1}}}+\sqrt{\frac{\mu}{b_{3}}} \frac{x\left(1+e_{1}\right)}{\sqrt{2-x^{2}\left(1+e_{1}\right)}}=0
$$

After some reductions, we obtain the value of $(x)_{\text {Min. }}$ on the form

$$
(x)_{M i n}= \pm \sqrt{\frac{2 b_{3}}{\left(1+e_{1}\right)\left(b_{1}+b_{3}\right)}}
$$

Finally

$$
\begin{gather*}
\left(\Delta v_{T}\right)_{M i n}=\left(\Delta v_{A}\right)_{M i n}+\left(\Delta v_{B}\right)_{M i n} \\
\left(\Delta v_{T}\right)_{M i n}=\sqrt{\frac{2 \mu b_{3}}{b_{1}\left(b_{1}+b_{3}\right)}}-\sqrt{\frac{\mu\left(1+e_{1}\right)}{b_{1}}}+\sqrt{\frac{\mu\left(1+e_{2}\right)}{b_{3}}}-\sqrt{\frac{2 \mu b_{1}}{b_{3}\left(b_{1}+b_{3}\right)}} \tag{28}
\end{gather*}
$$

Or explicitly in terms of the elements a, e:

$$
\begin{align*}
& \left(\Delta v_{T}\right)_{M i n}=\sqrt{\frac{2 \mu a_{2}\left(1-e_{2}\right)}{a_{1}\left(1-e_{1}\right)\left\{a_{1}\left(1-e_{1}\right)+a_{2}\left(1-e_{2}\right)\right\}}}-\sqrt{\frac{\mu\left(1+e_{1}\right)}{a_{1}\left(1-e_{1}\right)}} \\
& +\sqrt{\frac{\mu\left(1+e_{2}\right)}{a_{2}\left(1-e_{2}\right)}}-\sqrt{\frac{2 \mu a_{1}\left(1-e_{1}\right)}{a_{2}\left(1-e_{2}\right)\left\{a_{1}\left(1-e_{1}\right)+a_{2}\left(1-e_{2}\right)\right\}}} \tag{29}
\end{align*}
$$



Figure 3 Apogee of transfer orbit coincides with perigee of final orbit (Initial impulse at perigee)

For the third configuration (Fig.3), we have

$$
\begin{align*}
& a_{1}\left(1+e_{1}\right)=a_{T}\left(1+e_{T}\right) \\
& a_{T}\left(1-e_{T}\right)=a_{2}\left(1-e_{2}\right) \\
& v_{A 1}=\sqrt{\frac{\mu\left(1-e_{1}\right)}{b_{2}}}  \tag{30}\\
& v_{A 2}=\sqrt{\frac{\mu\left(1-e_{T}\right)}{b_{2}}}
\end{align*}
$$

Let

$$
x=\frac{v_{A 2}}{v_{A 1}}=\sqrt{\frac{1-e_{T}}{1-e_{1}}}>1
$$

i.e.

$$
\begin{gathered}
e_{T}=1-x^{2}\left(1-e_{1}\right) \\
a_{T}=\frac{a_{1}\left(1+e_{1}\right)}{1+e_{T}}=\frac{a_{2}\left(1-e_{2}\right)}{1-e_{T}}
\end{gathered}
$$

i.e.

$$
a_{T}=\frac{b_{2}}{1+e_{T}}=\frac{b_{3}}{1-e_{T}}
$$

$$
\begin{align*}
& a_{T}=\frac{b_{2}}{2-x^{2}\left(1-e_{1}\right)}=\frac{b_{3}}{x^{2}\left(1-e_{1}\right)} \\
& \Delta v_{A}=v_{A 2}-v_{A 1}=x v_{A 1}-v_{A 1}=(x-1) v_{A 1} \\
& \Delta v_{A}=(x-1) \sqrt{\frac{\mu\left(1-e_{1}\right)}{b_{2}}}  \tag{32}\\
& v_{B 1}=\sqrt{\frac{\mu\left(1+e_{T}\right)}{a_{T}\left(1-e_{T}\right)}} \\
& v_{B 2}=\sqrt{\frac{\mu\left(1+e_{2}\right)}{a_{2}\left(1-e_{2}\right)}} \\
& \Delta v_{B}=v_{B 2}-v_{B 1}=\sqrt{\frac{\mu\left(1+e_{2}\right)}{b_{3}}}-\sqrt{\frac{\mu\left\{2-x^{2}\left(1-e_{1}\right)\right\}}{b_{3}}} \tag{33}
\end{align*}
$$

Optimum condition is:

$$
\frac{d}{d x}\left(\Delta v_{T}\right)=\frac{d}{d x}\left(\Delta v_{A}\right)+\frac{d}{d x}\left(\Delta v_{B}\right)=0
$$

Then from Eqs. (32), (33) we get

$$
\sqrt{\frac{\mu\left(1-e_{1}\right)}{b_{2}}}+\sqrt{\frac{\mu}{b_{3}}} \frac{x\left(1-e_{1}\right)}{\sqrt{2-x^{2}\left(1-e_{1}\right)}}=0
$$

After some simple algebraic reductions, we get

$$
\begin{gather*}
(x)_{\text {Min }}= \pm \sqrt{\frac{2 b_{3}}{\left(1-e_{1}\right)\left(b_{2}+b_{3}\right)}} \\
\left(\Delta v_{T}\right)_{M i n}=\sqrt{\frac{2 \mu b_{3}}{b_{2}\left(b_{2}+b_{3}\right)}}-\sqrt{\frac{\mu\left(1-e_{1}\right)}{b_{2}}}+\sqrt{\frac{\mu\left(1+e_{2}\right)}{b_{3}}}-\sqrt{\frac{2 \mu b_{2}}{b_{3}\left(b_{2}+b_{3}\right)}} \tag{34}
\end{gather*}
$$

Or explicitly,

$$
\begin{align*}
& \left(\Delta v_{T}\right)_{M i n}=\sqrt{\frac{2 \mu a_{2}\left(1-e_{2}\right)}{a_{1}\left(1+e_{1}\right)\left\{a_{1}\left(1+e_{1}\right)+a_{2}\left(1-e_{2}\right)\right\}}}-\sqrt{\frac{\mu\left(1-e_{1}\right)}{a_{1}\left(1+e_{1}\right)}} \\
& +\sqrt{\frac{\mu\left(1+e_{2}\right)}{a_{2}\left(1-e_{2}\right)}}-\sqrt{\frac{2 \mu a_{1}\left(1+e_{1}\right)}{a_{2}\left(1-e_{2}\right)\left\{a_{1}\left(1+e_{1}\right)+a_{2}\left(1-e_{2}\right)\right\}}} \tag{35}
\end{align*}
$$

For the fourth configuration (Fig. 4), we have

$$
\begin{aligned}
a_{1}\left(1+e_{1}\right)=a_{T}\left(1-e_{T}\right) & a_{T}\left(1+e_{T}\right)=a_{2}\left(1+e_{2}\right) \\
v_{A 1}=\sqrt{\frac{\mu\left(1-e_{1}\right)}{b_{2}}} & v_{A 2}=\sqrt{\frac{\mu\left(1+e_{T}\right)}{b_{2}}}
\end{aligned}
$$



Figure 4 Apogee of transfer orbit coincides with apogee of final orbit (Initial impulse at apogee)

$$
x=\frac{v_{A 2}}{v_{A 1}}=\sqrt{\frac{1+e_{T}}{1-e_{1}}}>1
$$

i.e.

$$
\begin{align*}
& e_{T}=x^{2}\left(1-e_{1}\right)-1  \tag{36}\\
& a_{T}=\frac{a_{1}\left(1+e_{1}\right)}{1-e_{T}}=\frac{a_{2}\left(1+e_{2}\right)}{1+e_{T}} \\
& a_{T}=\frac{b_{2}}{2-x^{2}\left(1-e_{1}\right)}=\frac{b_{4}}{x^{2}\left(1-e_{1}\right)} \\
& \Delta v_{A}=v_{A 2}-v_{A 1}=x v_{A 1}-v_{A 1}=(x-1) v_{A 1} \\
& \Delta v_{A}=(x-1) \sqrt{\frac{\mu\left(1-e_{1}\right)}{b_{2}}}  \tag{37}\\
& v_{B 1}=\sqrt{\frac{\mu\left(1-e_{T}\right)}{b_{4}}} \\
& v_{B 2}=\sqrt{\frac{\mu\left(1-e_{2}\right)}{b_{4}}} \\
& \Delta v_{B}=v_{B 2}-v_{B 1}=\sqrt{\frac{\mu\left(1-e_{2}\right)}{b_{4}}}-\sqrt{\frac{\mu\left\{2-x^{2}\left(1-e_{1}\right)\right\}}{b_{4}}}  \tag{38}\\
& \Delta v_{T}=\Delta v_{A}+\Delta v_{B}
\end{align*}
$$

Optimum condition is

$$
\frac{d}{d x}\left(\Delta v_{T}\right)=\frac{d}{d x}\left(\Delta v_{A}\right)+\frac{d}{d x}\left(\Delta v_{B}\right)=0
$$

Then from Eqs. (37), (38) we get

$$
\begin{aligned}
& \sqrt{\frac{\mu\left(1-e_{1}\right)}{b_{2}}}+\sqrt{\frac{\mu}{b_{4}}} \frac{x\left(1-e_{1}\right)}{\sqrt{2-x^{2}\left(1-e_{1}\right)}}=0 \\
& (x)_{\text {Min }}= \pm \sqrt{\frac{2 b_{4}}{\left(1-e_{1}\right)\left(b_{2}+b_{4}\right)}} \\
& \left(\Delta v_{T}\right)_{M i n}=\left(\Delta v_{A}\right)_{M i n}+\left(\Delta v_{B}\right)_{M i n}
\end{aligned}
$$

Finally,

$$
\begin{align*}
& \left(\Delta v_{T}\right)_{M i n}=\sqrt{\frac{2 \mu b_{4}}{b_{2}\left(b_{2}+b_{4}\right)}}-\sqrt{\frac{\mu\left(1-e_{1}\right)}{b_{2}}} \\
& +\sqrt{\frac{\mu\left(1-e_{2}\right)}{b_{4}}}-\sqrt{\frac{2 \mu b_{2}}{b_{4}\left(b_{2}+b_{4}\right)}} \tag{39}
\end{align*}
$$

Or explicitly

$$
\begin{align*}
& \left(\Delta v_{T}\right)_{M i n}=\sqrt{\frac{2 \mu a_{2}\left(1+e_{2}\right)}{a_{1}\left(1+e_{1}\right)\left\{a_{1}\left(1+e_{1}\right)+a_{2}\left(1+e_{2}\right)\right\}}}-\sqrt{\frac{\mu\left(1-e_{1}\right)}{a_{1}\left(1+e_{1}\right)}} \\
& +\sqrt{\frac{\mu\left(1-e_{2}\right)}{a_{2}\left(1+e_{2}\right)}}-\sqrt{\frac{2 \mu a_{1}\left(1+e_{1}\right)}{a_{2}\left(1+e_{2}\right)\left\{a_{1}\left(1+e_{1}\right)+a_{2}\left(1+e_{2}\right)\right\}}} \tag{40}
\end{align*}
$$

## 3. Numerical calculations

We consider the Earth - Mars Hohmann elliptic transfer to perform an approximative check for the validity of the above calculations. We have
$\mathrm{a}_{1}=1 \mathrm{AU} ; \mathrm{a}_{2}=1.5237 \mathrm{AU} ; \mathrm{e}_{1}=0.0167 ; \mathrm{e}_{2}=0.0934$
where subscript 1 refers to the Earth and subscript 2 refers to the Mars.
We have the following table for the four configurations:

| Fig. | $\left(\mathrm{a}_{T}\right)_{\text {Min }}$ | $\left(\mathrm{e}_{T}\right)_{\text {Min }}$ | $(\mathrm{x})_{\text {Min }}$ | $\left(\Delta \mathrm{v}_{T}\right)_{\text {Min }}$ | $\mathrm{a}_{2}$ |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 1 | 1.3247 | 0.2577 | 1.1122 | 0.1843 | 1.5237 |
| 2 | 1.1823 | 0.1683 | 1.0720 | 0.1870 | 1.5236 |
| 3 | 1.1990 | -0.1521 | 1.0824 | 0.1873 | 1.5237 |
| 4 | 1.3414 | 0.2420 | 1.1239 | 0.1850 | 1.5237 |

We note that $\left(e_{T}\right)_{\text {Min }}$ for Fig. 3. is negative value, according to the sketch of this figure and the Eq. (30).

We assume that $a_{1}, a_{2}$ the semi-major axes of the Earth and Mars are equal to the mean distances of the two planets from the primary (the Sun). Evidently $\left(\Delta V_{T}\right)_{M i n}$ of Fig. 1. is the most economic.

## 4. Concluding Remarks

The choice of x as our variable leads to the most simple and exact formulate of the problem. After the resolution of the second degree equation in x arising from the optimum condition, we can determine the unique values $\left(e_{T}\right)_{M i n},\left(a_{T}\right)_{M i n}$ from Eqs. (9), (10), (25),(31),(36) knowing the given values of $a_{1}, e_{1}, a_{2}, e_{2}$ of the initial and final orbit. The minimum characteristic velocity $\left(\Delta v_{T}\right)_{\text {Min }}$ Eqs. (22), (28), (34), (39) are obviously expressed in terms of the initial and final orbital elements (the major axes and the eccentricities $a_{1}, a_{2}, e_{1}, e_{2}$ ). The optimization procedure is based on formulas stemming from first principles considerations. It is not a special case arising from the general problem, when we assume non coplanar trajectories. We verified the correctness of the approach by the assignment of the approximative value of $a_{2}$ (the semi major axis of the final orbit), from the formula
$2 a_{2}=\left\{a_{T}\left(1+e_{T}\right)+a_{2}\left(1-e_{2}\right)\right\}$, for Fig. 1 and 4,
$2 a_{2}=\left\{a_{T}\left(1+e_{T}\right)+a_{2}\left(1+e_{2}\right)\right\}$, for Fig. 2
$2 a_{2}=\left\{a_{T}\left(1-e_{T}\right)+a_{2}\left(1+e_{2}\right)\right\}$, for Fig. 3.
In this paper, we consider four feasible configurations for this transfer problem, two of them are relevant to the peri-apse perpendicular initial impulse (Fig. 1. and Fig. 2.), the other two are relevant to initial perpendicular apo-apse impulse (Fig. 3. and Fig. 4.).

This approach is new, elementary, and straightforward. It avoids many complexities that appear in other works, thus it is advantageous for this particular transfer problem, and it is a proof that the generalized Hohmann transfer is itself a minimum orbit transfer system.

## References

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## Nomenclature

$x \quad$ ratio of velocities after and before initial impulse.
$a_{1} \quad$ semi-major axis of initial orbit.
$a_{2} \quad$ semi-major axis of final orbit.
$e_{1} \quad$ eccentricity of initial orbit.
$e_{2} \quad$ eccentricity of final orbit.
$a_{T} \quad$ semi-major axis of transfer orbit.
$e_{T} \quad$ eccentricity of transfer orbit.
$v_{A 1} \quad$ peri-apse velocity in initial orbit at point A.
$v_{A 2}$ peri-apse velocity of transfer orbit at point A.
$v_{B 1} \quad$ apo-apse velocity of transfer orbit at point $B$.
$v_{B 2}$ apo-apse velocity in final orbit at point B .
$\Delta v_{1} \quad$ increment of velocity at A.
$\Delta v_{2} \quad$ increment of velocity at B.
$\Delta v_{T} \quad$ characteristic velocity.
$\mu \quad$ constant of gravitation.

